



Extinguishing Poisson's Spot with Linear Programming

Robert J. Vanderbei

2008 May 13

SIAG/OPT 2008
Boston, MA

<http://www.princeton.edu/~rvdb>

Are We Alone?



Indirect Detection Methods

Almost 300 planets found so far

Wobble Methods

Radial Velocity.

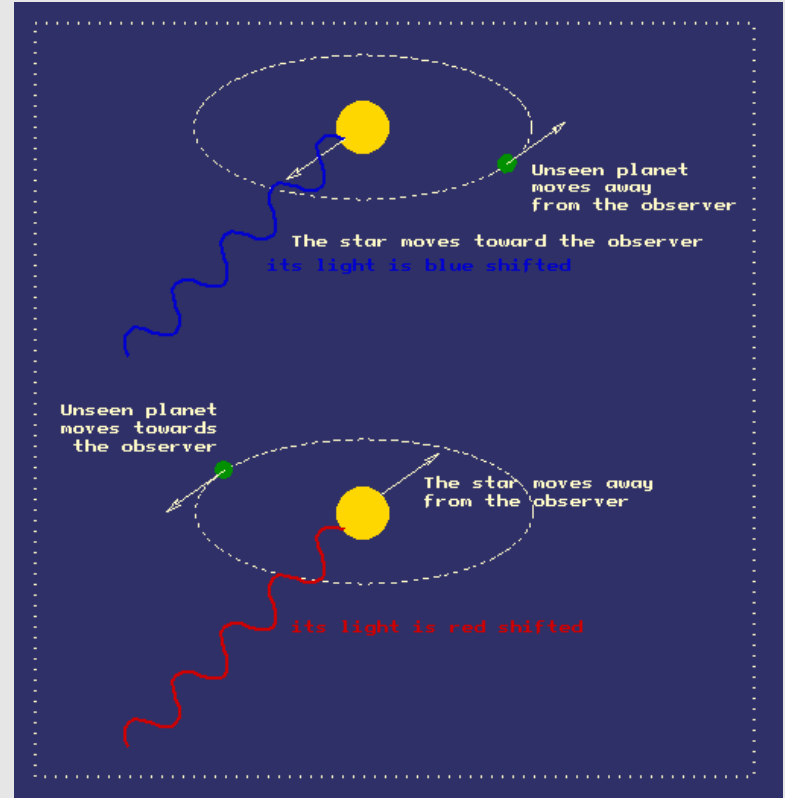
For edge-on systems.

Measure periodic doppler shift.

Astrometry.

Best for face-on systems.

Measure circular wobble against background stars.

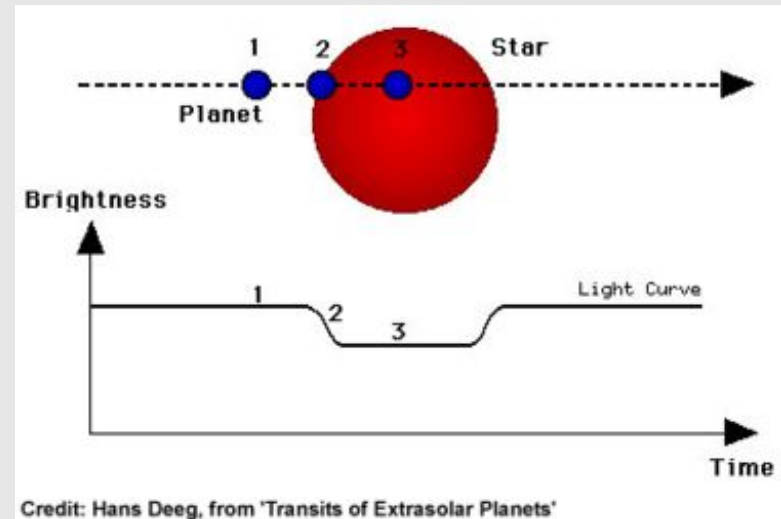


Transit Method

- HD209458b confirmed both via RV and transit.
- Period: 3.5 days
- Separation: 0.045 AU (0.001 arcsecs)
- Radius: $1.3R_J$
- Intensity Dip: $\sim 1.7\%$
- Venus Dip = 0.01%, Jupiter Dip: 1%



Venus Transit (R.J. Vanderbei)

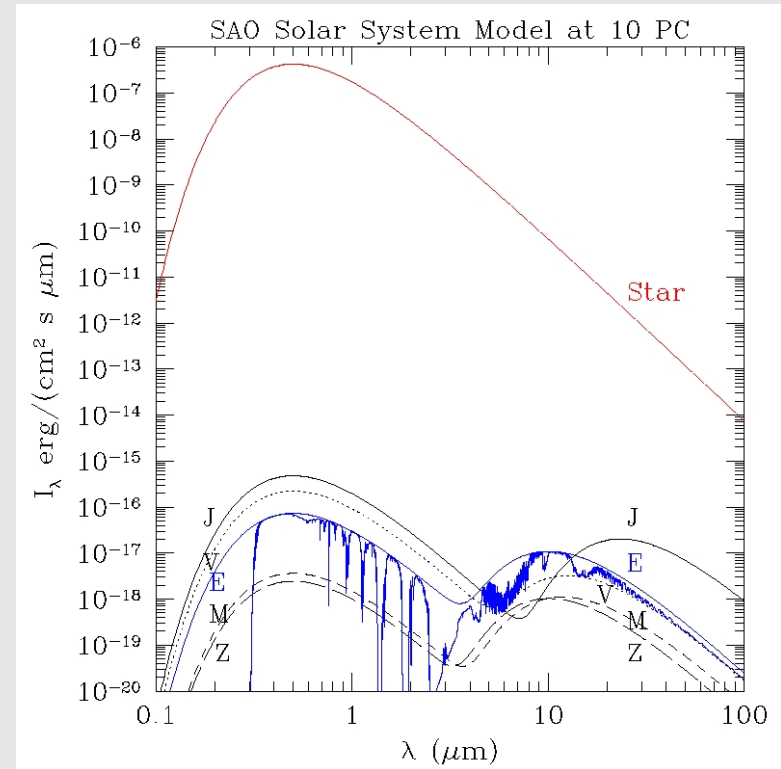


Credit: Hans Deeg, from 'Transits of Extrasolar Planets'

Direct Detection

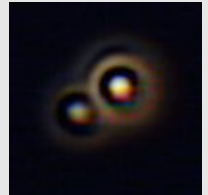
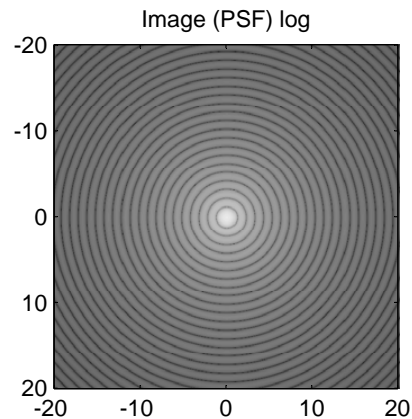
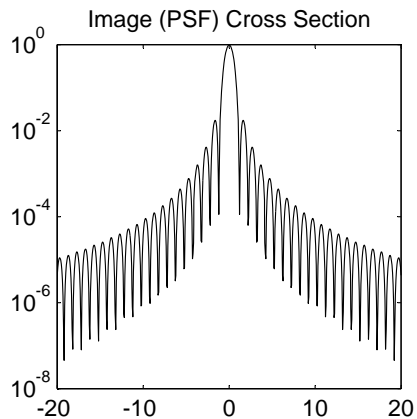
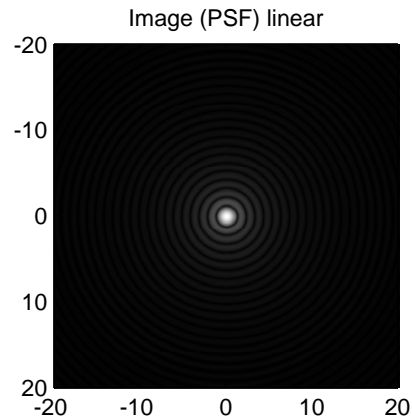
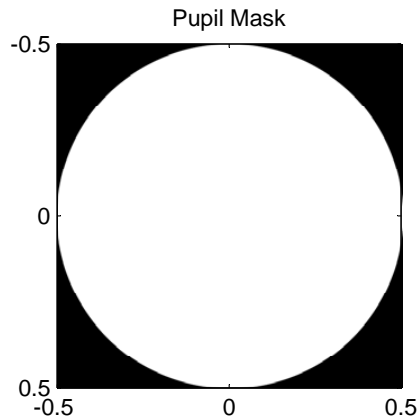
Why It's Hard

- *Bright Star/Faint Planet:* In visible light, our Sun is 10^{10} times brighter than Earth. That's 25 mags.
- *Close to Each Other:* A planet at 1 AU from a star at 10 parsecs can appear at most 0.1 arcseconds in separation.
- *Far from Us:* There are less than 100 Sun-like stars within 10 parsecs.



Telescope w/ Unobstructed Aperture

Doesn't Work! Requires an aperture measured in kilometers to mitigate diffraction effects.



Space-based Occulter (TPF-O)

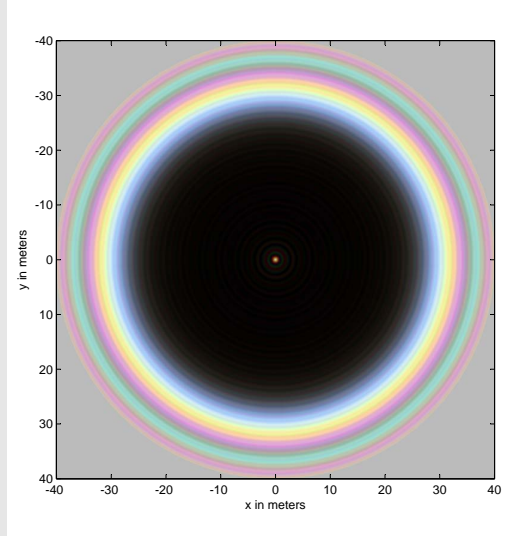
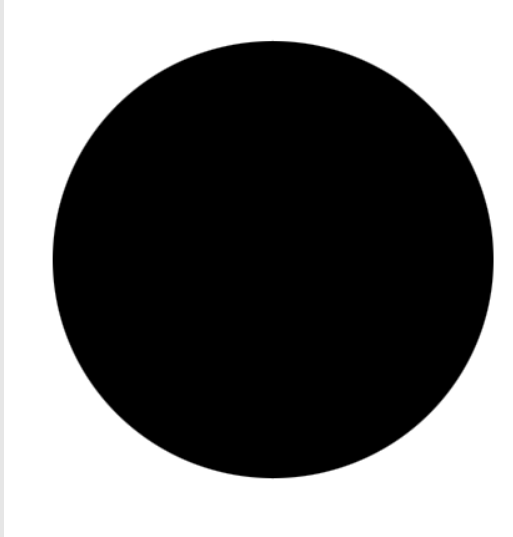


Telescope Aperture: 4m, Occulter Diameter: 50m, Occulter Distance: 72,000km

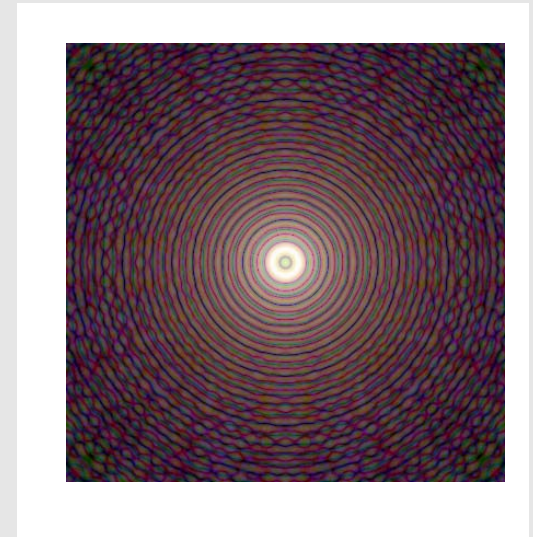
Plain External Occulter (Doesn't Work!)

Shadow \Rightarrow

Circular Occulter

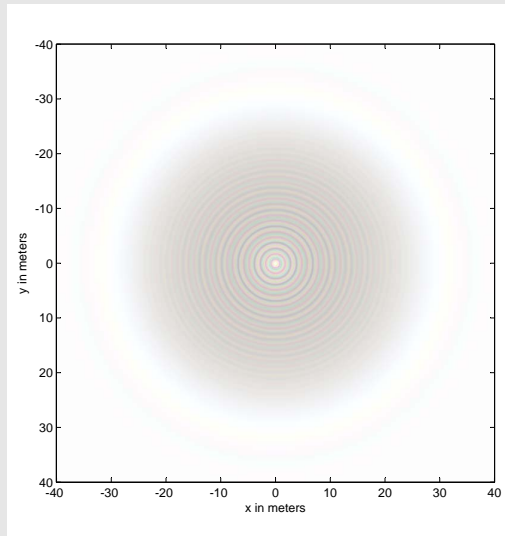


\Leftarrow Note bright spot at center
(Poisson's spot)



Telescope Image

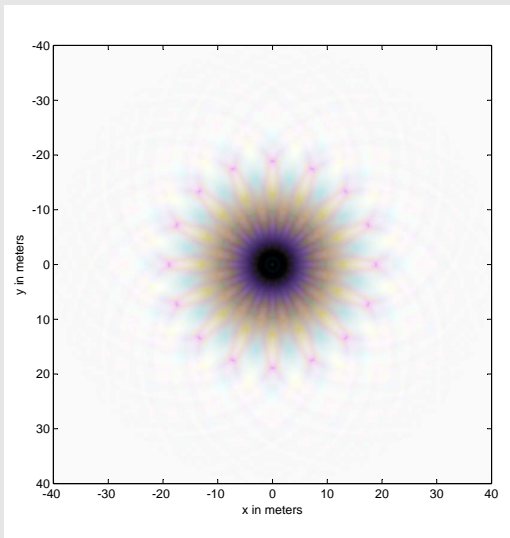
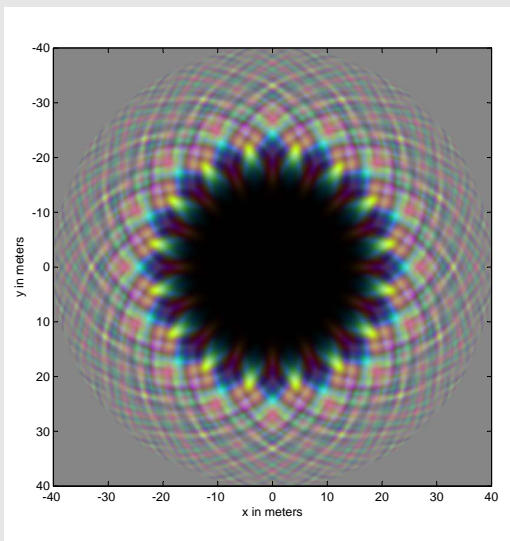
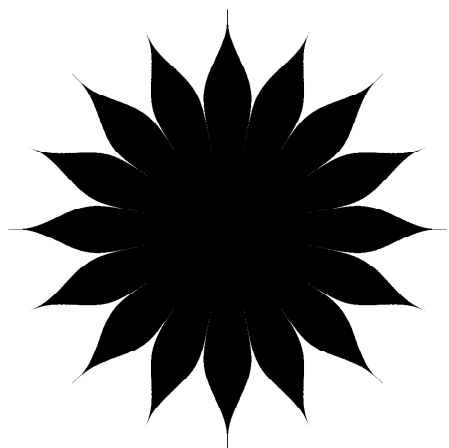
\Leftarrow Shadow (Log Stretch)



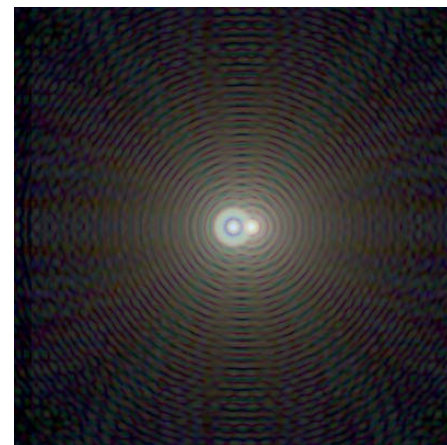
Shaped Occulter—Eliminates Poisson's Spot

Shadow \Rightarrow

Shaped Occulter



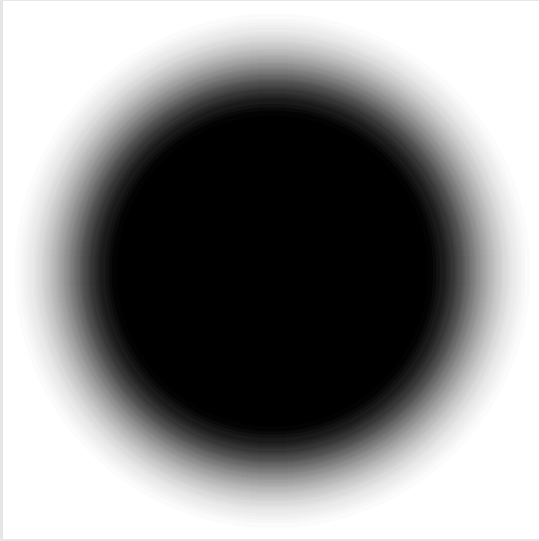
\Leftarrow Bright spot is gone



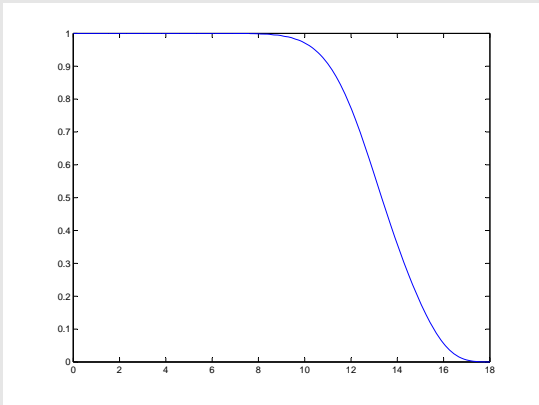
\Uparrow
Telescope image shows planet

\Leftarrow Shadow is dark
(Log Stretch)

Apodized Occulters



Apodized Occulter



Radial Attenuation $A(r)$

- The problem is *diffraction*.
- Abrupt edges create unwanted diffraction.
- *Solution*: Soften the edges with a partially transmitting material—an *apodizer*.
- Let $A(r, \theta)$ denote *attenuation* at location (r, θ) on the occulter.
- The *intensity* of the downstream light is given by the *square of the magnitude of the electric field* $E(\rho, \phi)$.
- *Babinet's principle* plus *Fresnel propagation* gives a formula for the downstream electric field:

$$E(\rho, \phi) = 1 - \frac{1}{i\lambda z} \int_0^\infty \int_0^{2\pi} e^{\frac{i\pi}{\lambda z}(r^2 + \rho^2 - 2r\rho \cos(\theta - \phi))} A(r, \theta) r d\theta dr.$$

where

- z is distance “downstream” and
- λ is wavelength of light.

Attenuation Profile Optimization

$$\begin{array}{ll} \text{minimize} & \gamma \\ \text{subject to} & -\gamma \leq \Re(E(\rho)) \leq \gamma \quad \text{for } \rho \in \mathcal{R}, \lambda \in \mathcal{L} \\ & -\gamma \leq \Im(E(\rho)) \leq \gamma \quad \text{for } \rho \in \mathcal{R}, \lambda \in \mathcal{L} \\ & A'(r) \leq 0 \quad \text{for } 0 \leq r \leq R \\ & -d \leq A''(r) \leq d \quad \text{for } 0 \leq r \leq R \end{array}$$

Specific choice:

$$R = 25, \quad d = 0.04, \quad \mathcal{R} = [0, 3], \quad \mathcal{L} = [0.4, 1.1] \times 10^{-6}$$

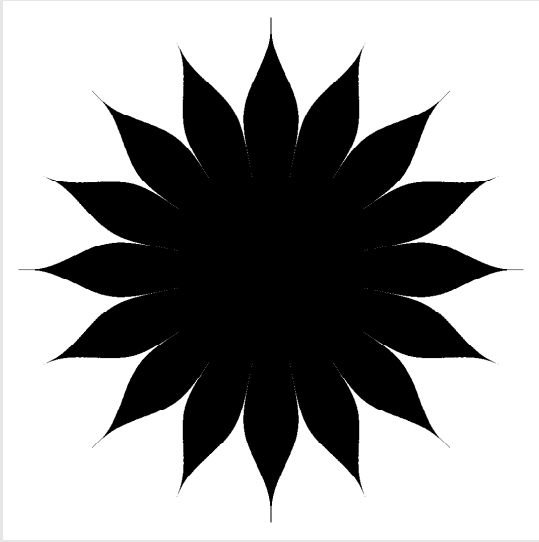
where all metric quantities are in meters.

An infinite dimensional linear programming problem.

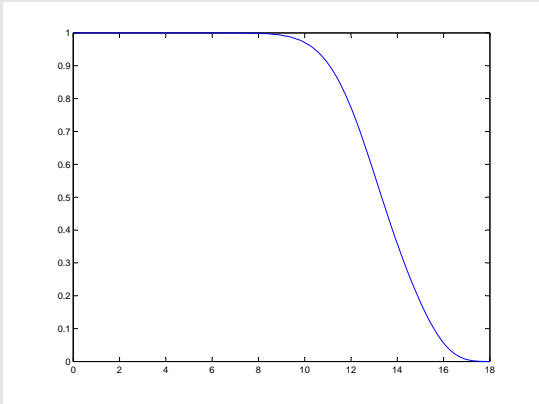
Discretize:

- $[0, R]$ into 5000 evenly space points.
- \mathcal{R} into 150 evenly spaced points.
- \mathcal{L} into increments of 0.1×10^{-6} .

Petal-Shaped Occulters



16-Petal Occulter $A(r, \theta)$



Radial Attenuation $A(r)$

- From Jacobi-Anger expansion we get:

$$E(\rho, \phi) = 1 - \frac{2\pi}{i\lambda z} \int_0^R e^{\frac{i\pi}{\lambda z}(r^2 + \rho^2)} J_0\left(\frac{2\pi r \rho}{\lambda z}\right) A(r) r dr \\ - \sum_{k=1}^{\infty} \frac{2\pi(-1)^k}{i\lambda z} \left(\int_0^R e^{\frac{i\pi}{\lambda z}(r^2 + \rho^2)} J_{kN}\left(\frac{2\pi r \rho}{\lambda z}\right) \frac{\sin(\pi k A(r))}{\pi k} r dr \right) \\ \times \left(2 \cos(kN(\phi - \frac{\pi}{2})) \right)$$

where N is the number of petals.

- For small ρ , truncated summation well-approximates full sum.
- Truncated after 10 terms.
- $\lambda \in [0.4, 1.1]$ microns.
- $z = 72,000$ km, $R = 25$ m.
- In angular terms, $R/z = 73$ mas.