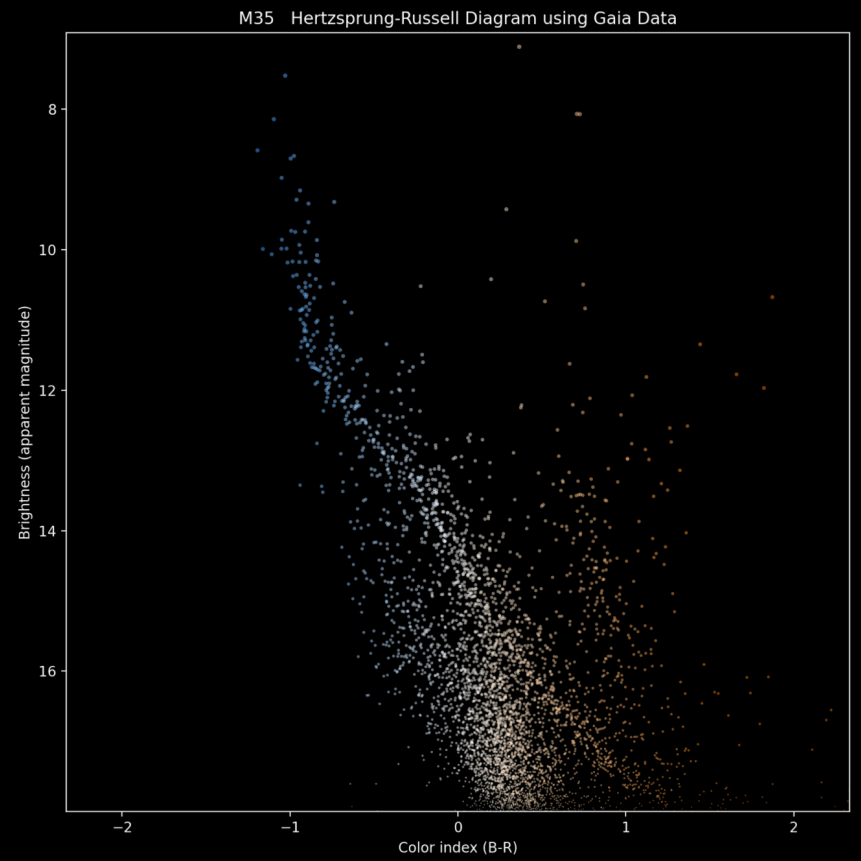


# HR-Diagrams of Open Clusters – Side-by-Side

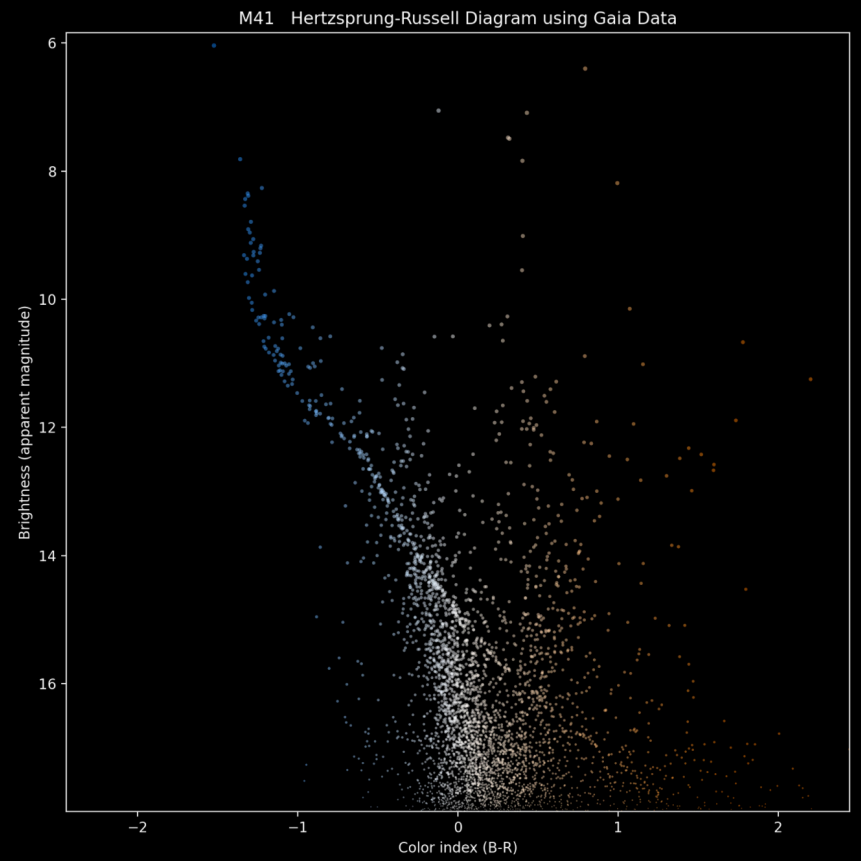
# HR-Diagrams of Open Clusters – Side-by-Side

Using Gaia Data

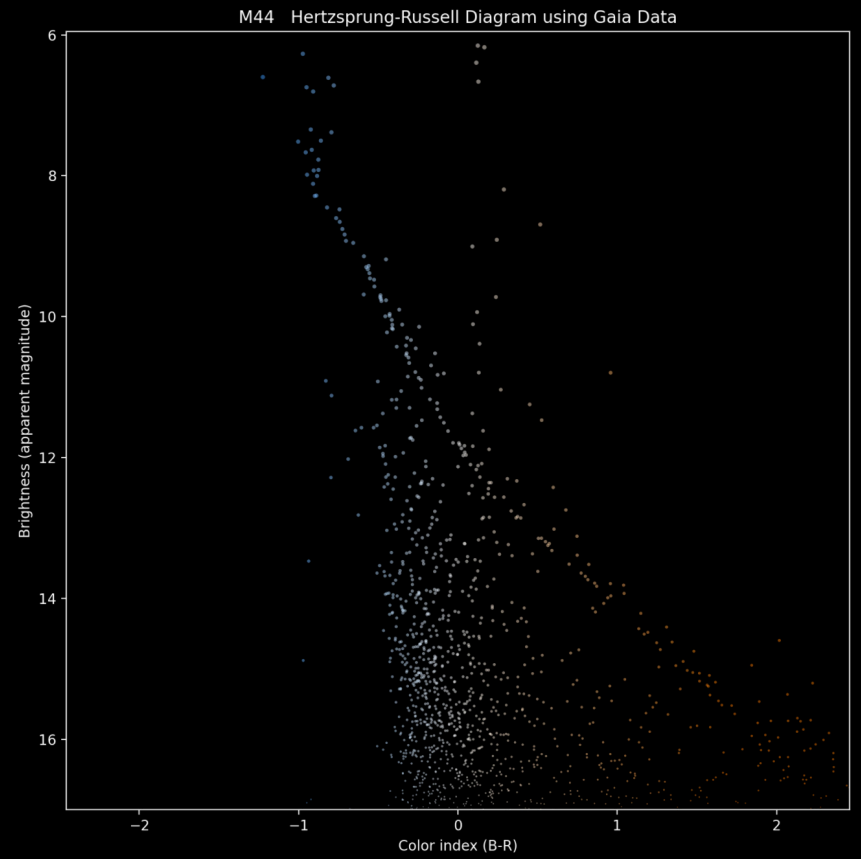
# M35



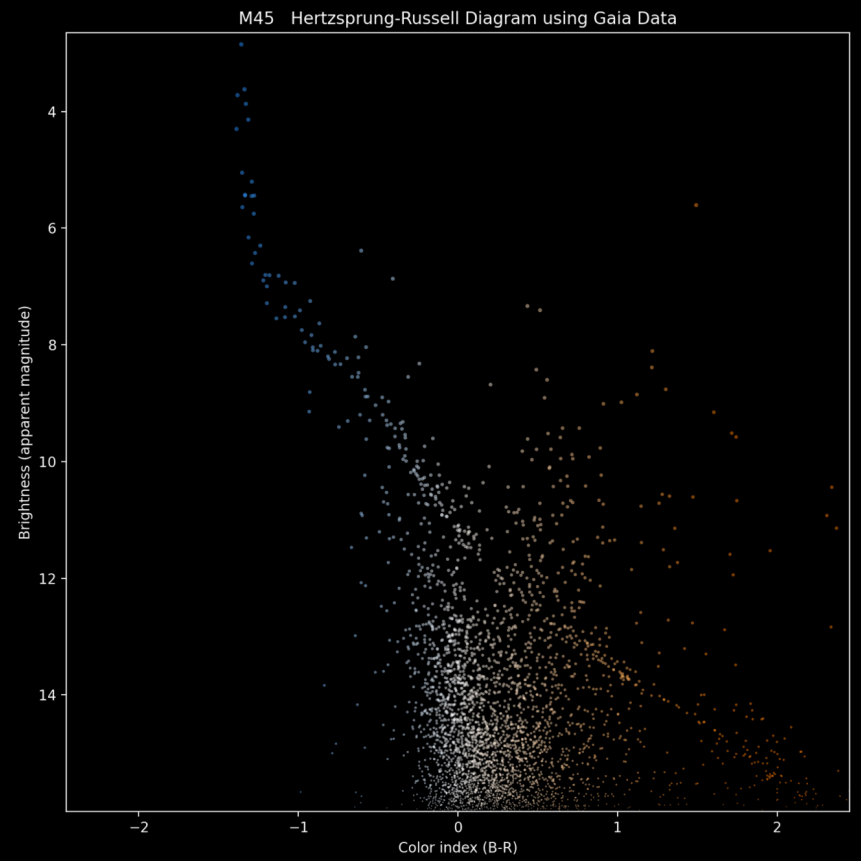
# M41



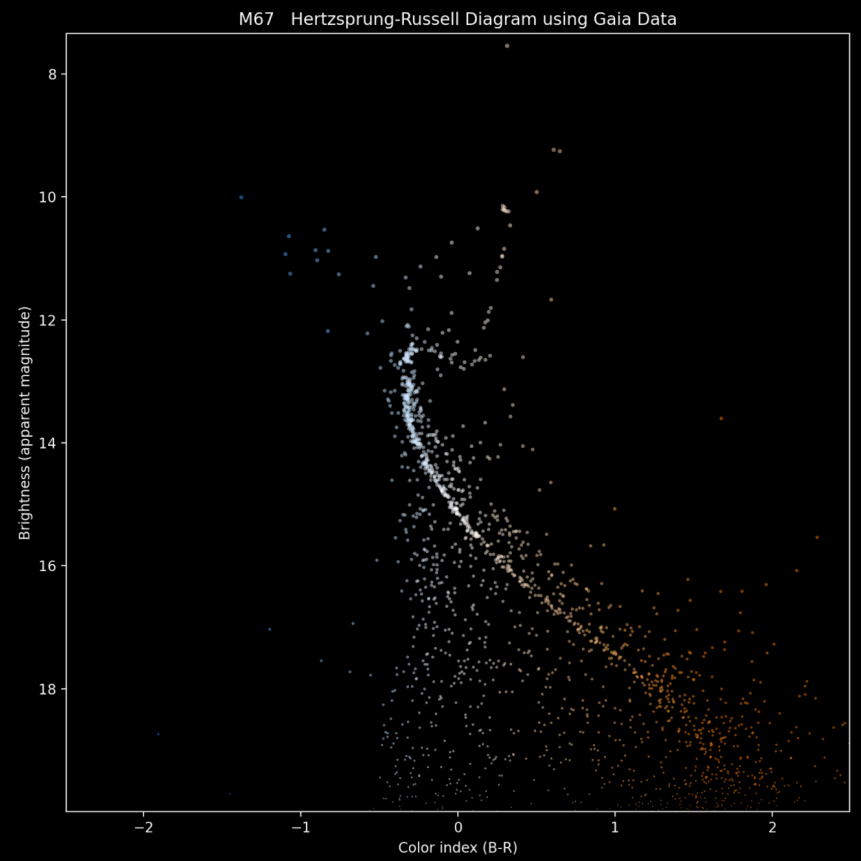
# M44



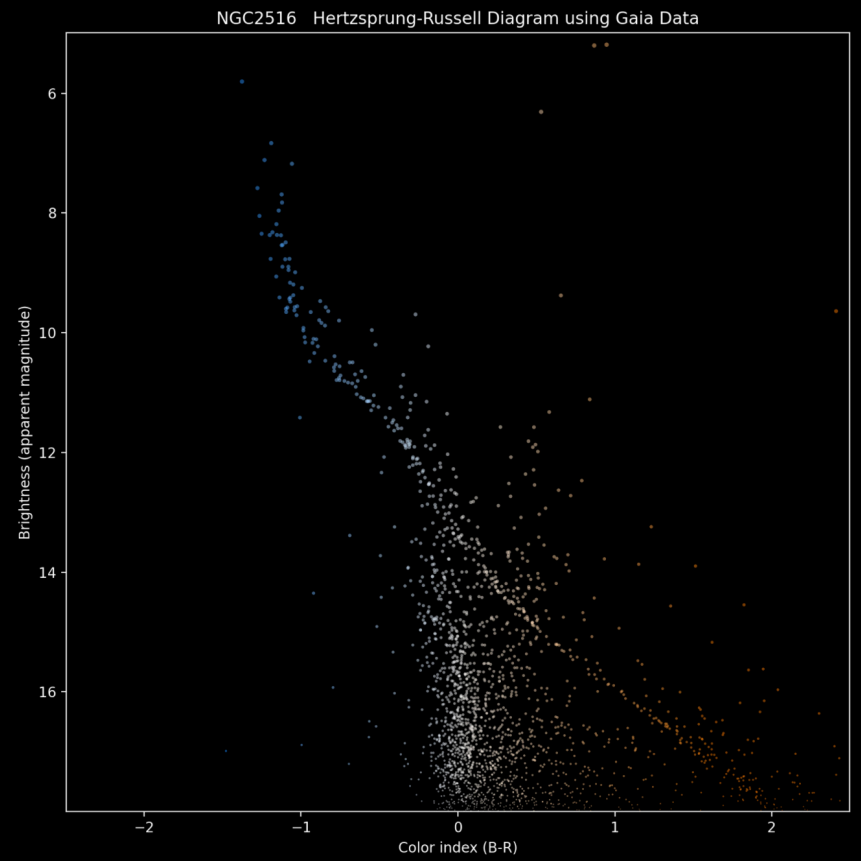
# M45



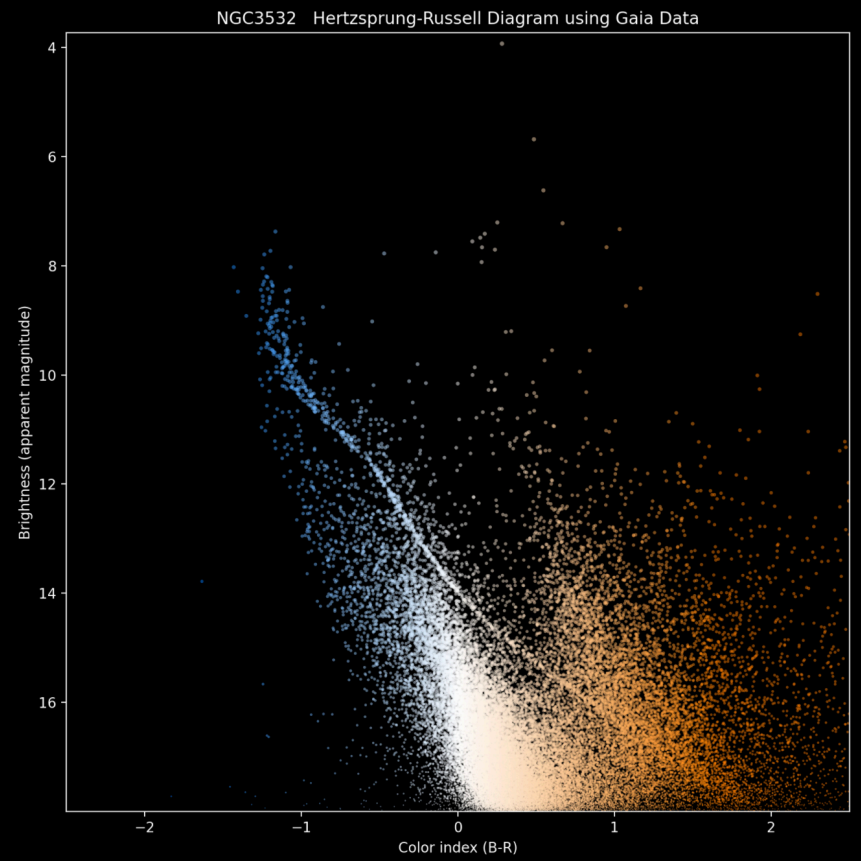
# M67



# NGC2516



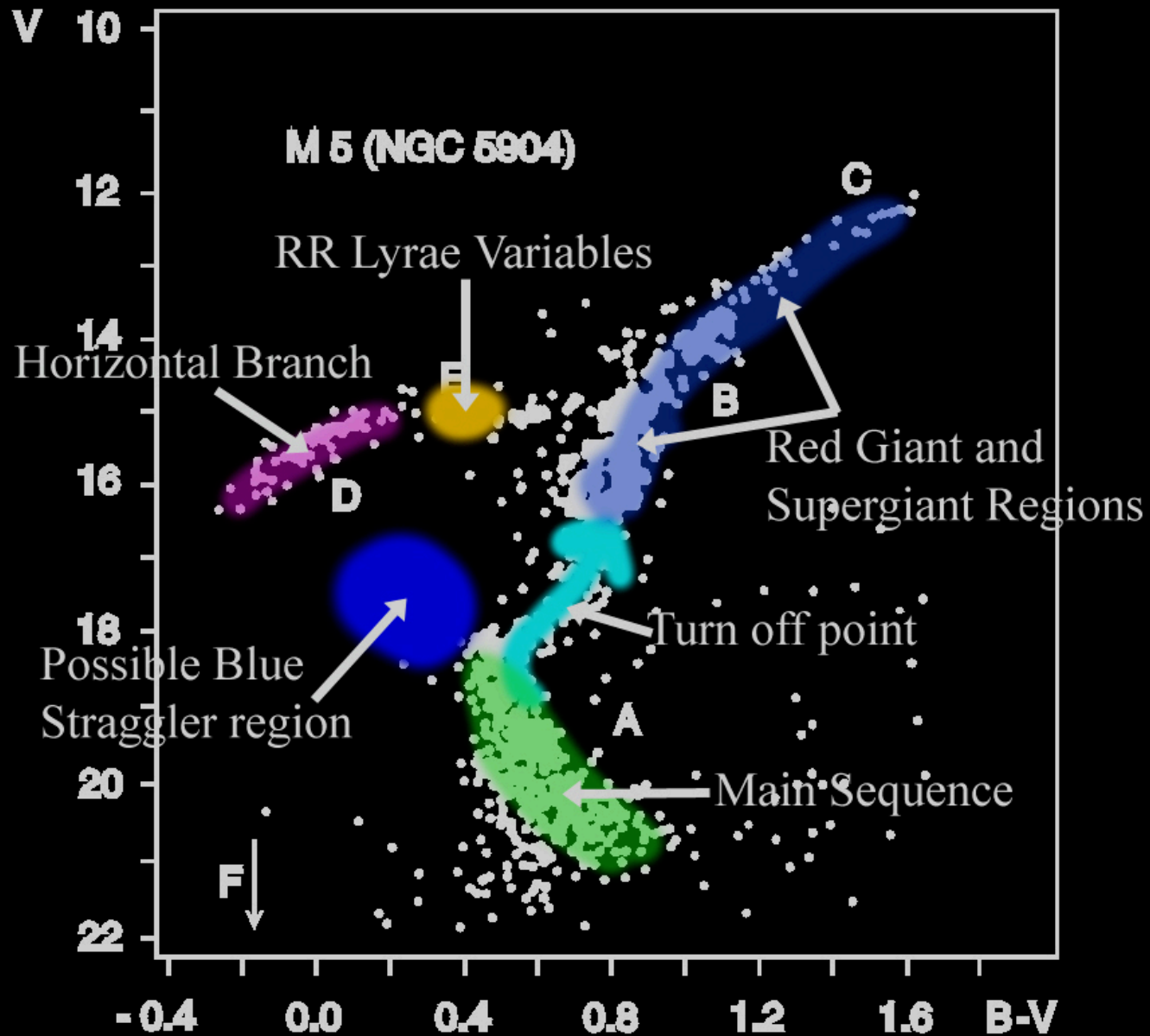
# NGC3532



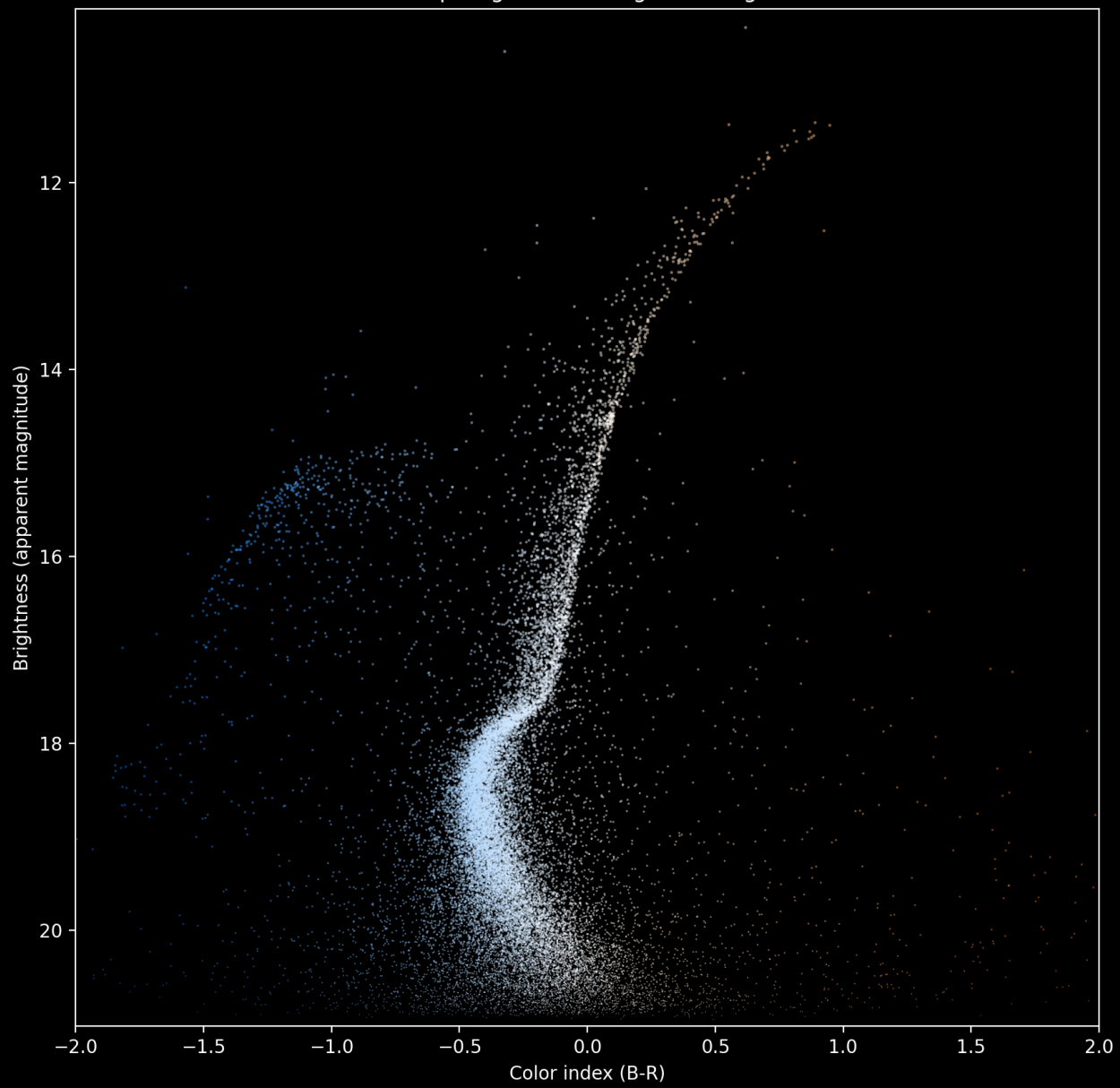
Any Questions?

# Variable Stars

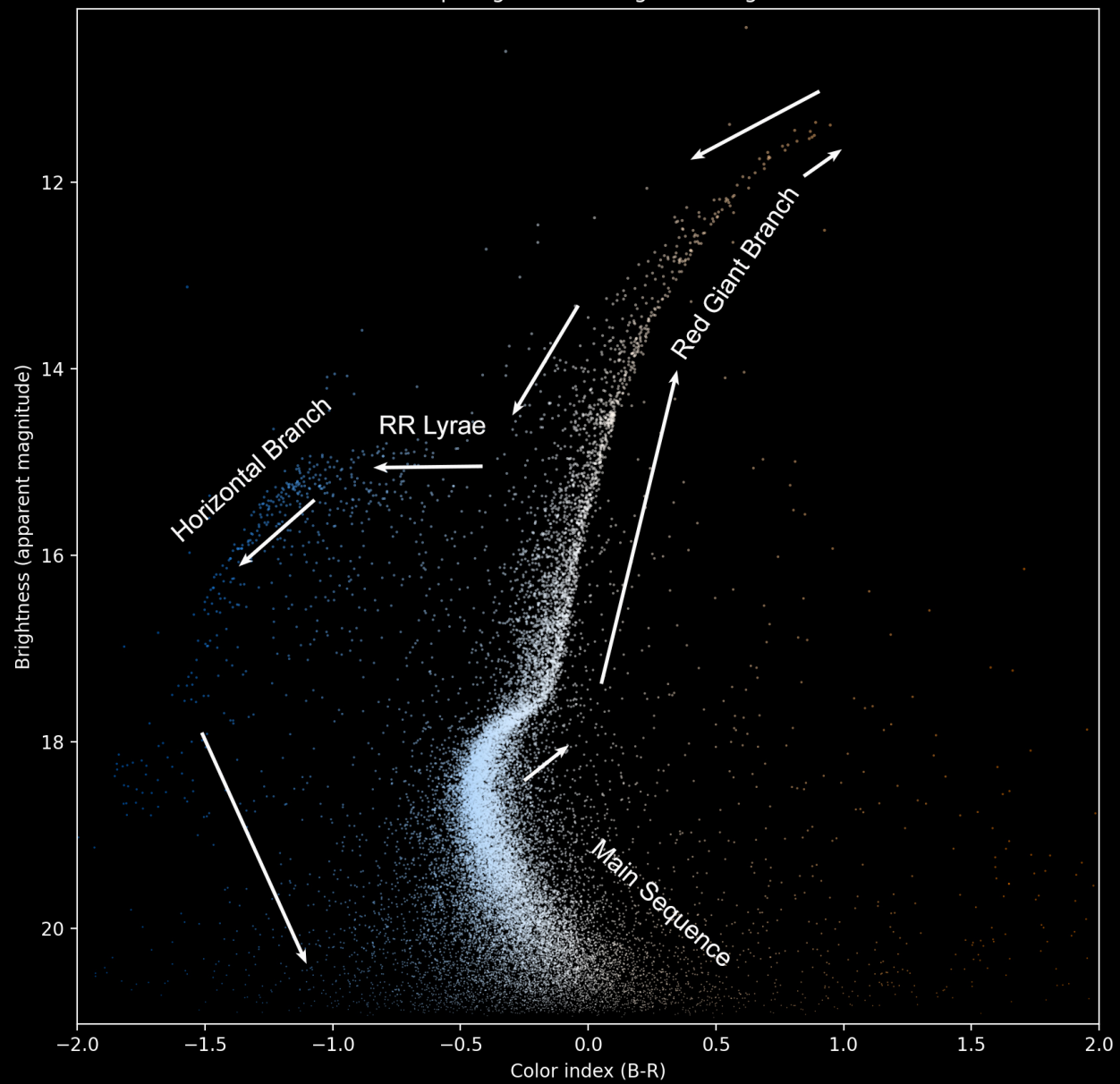
Adapted from SEDS (<http://www.seds.org>)



M13 Hertzprung-Russell Diagram using Gaia Data



M13 Hertzprung-Russell Diagram using Gaia Data

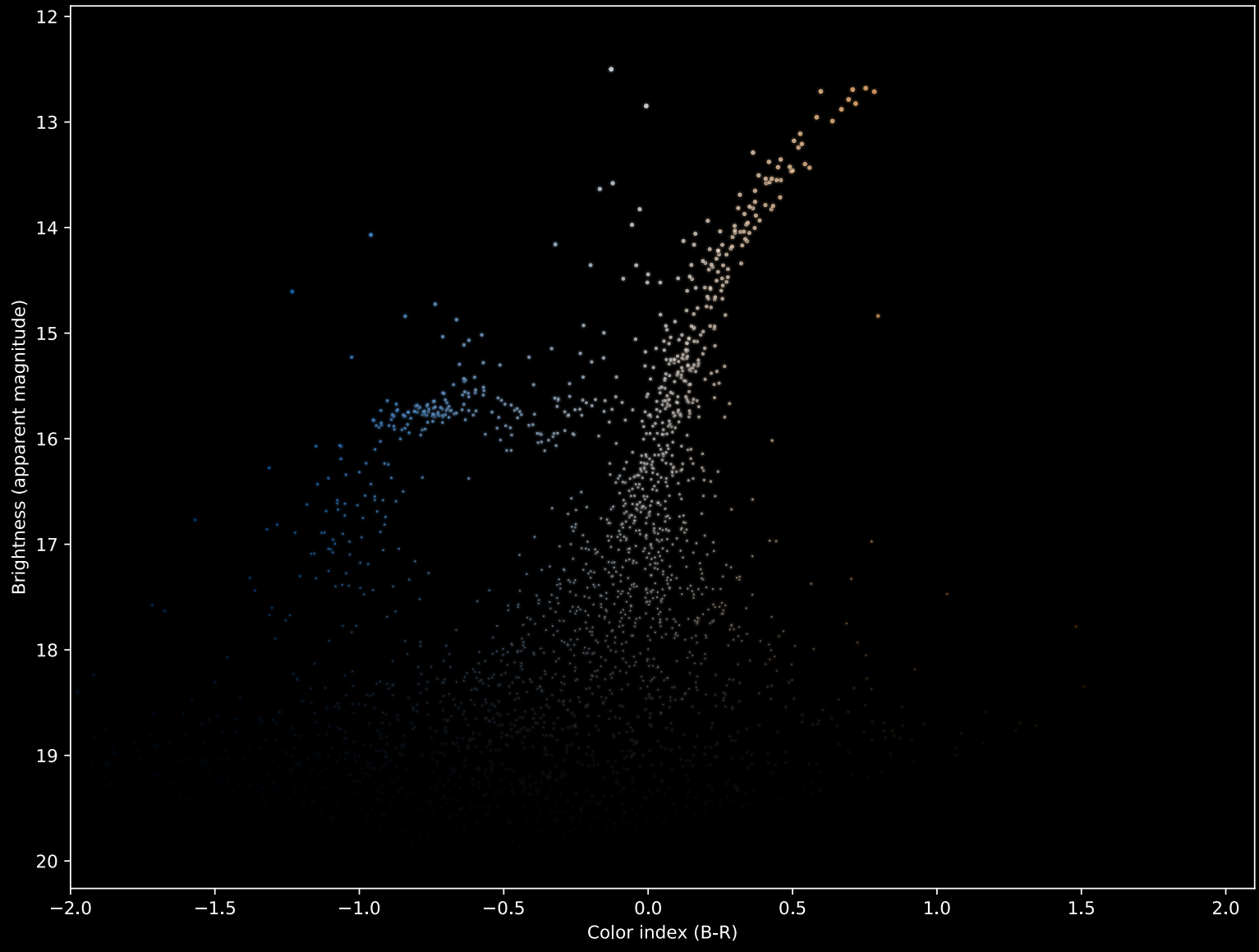


# Globular Cluster M15

[Click to see it](#)

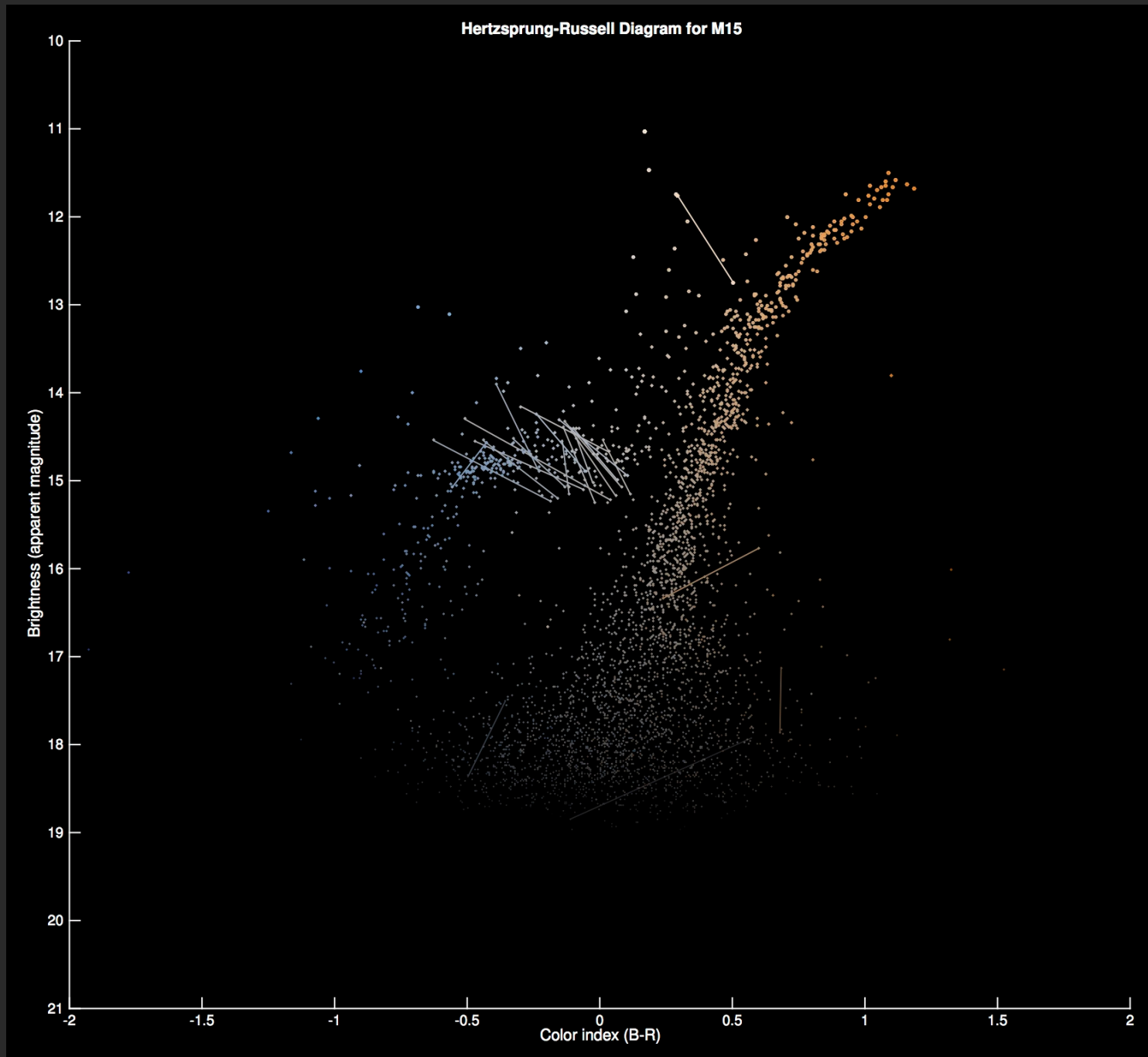


Hertzsprung-Russell Diagram for M15

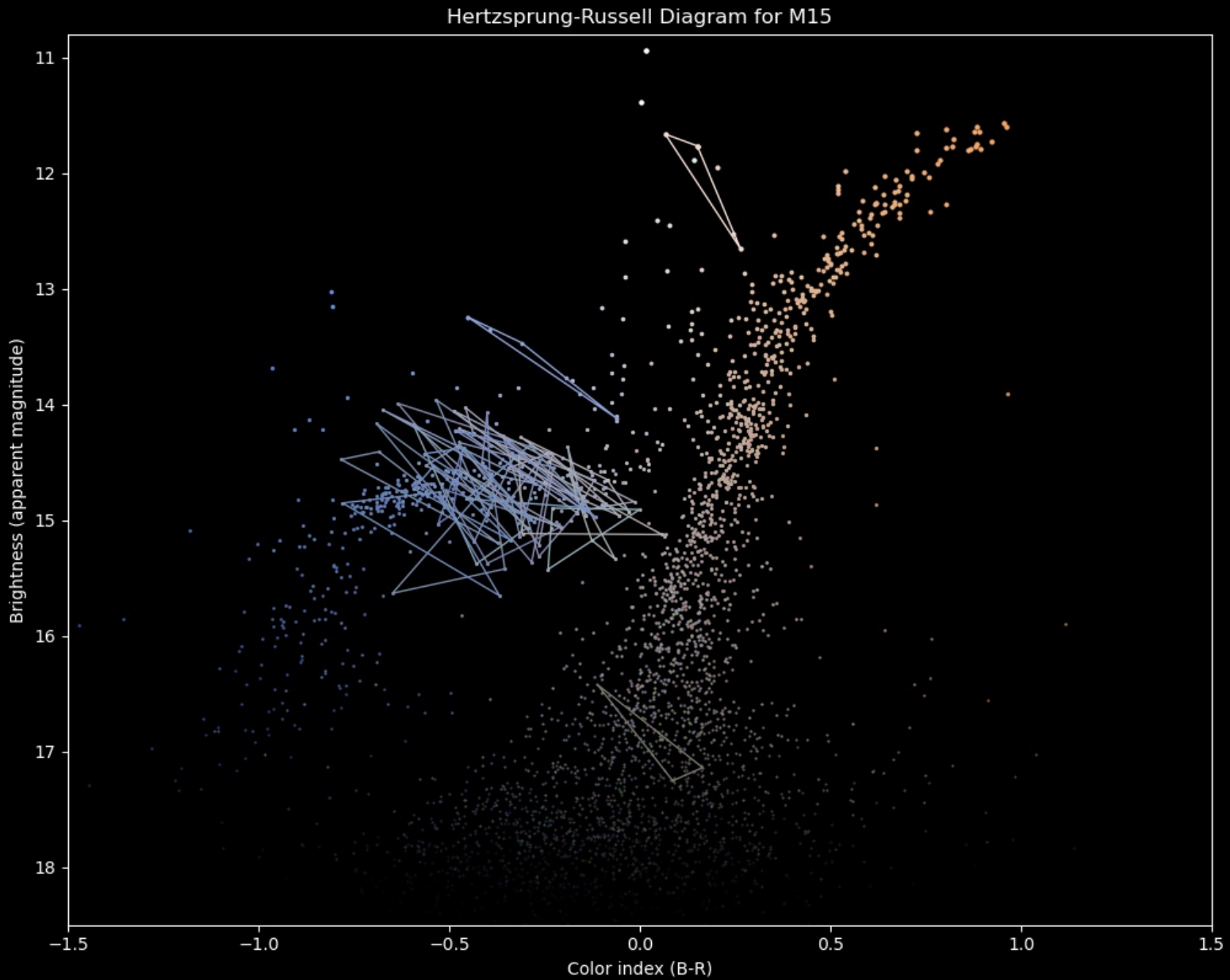


# HR-Diagram Showing RR-Lyrae Stars

[Click to see it](#)



# HR-Diagram Showing RR-Lyrae Stars



RR-Lyrae

distance: 258 parsecs

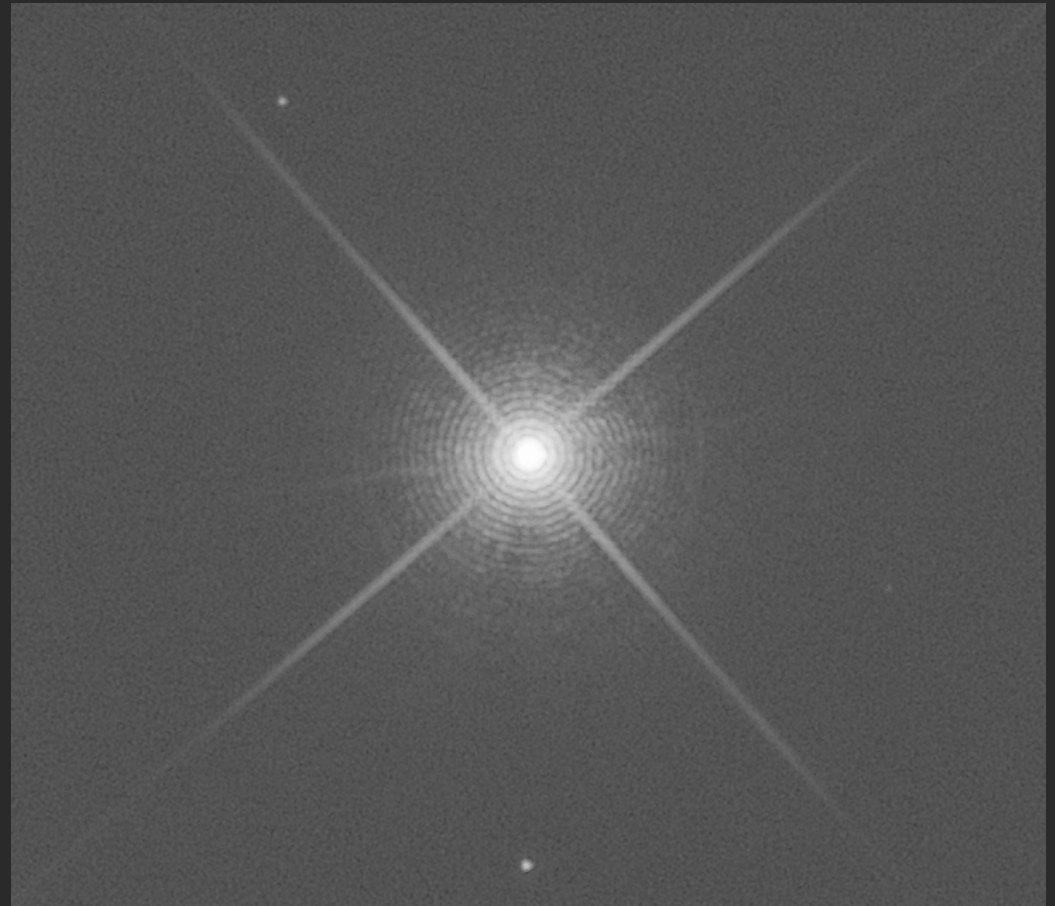
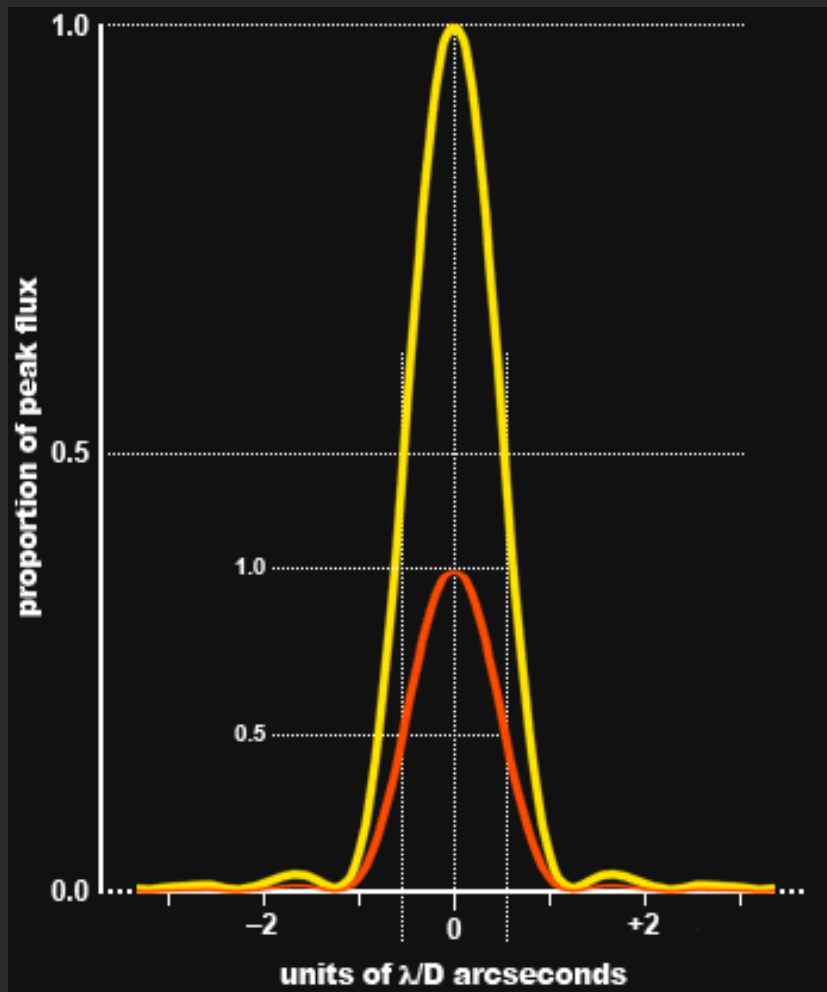


RR-Lyrae

distance: 258 parsecs



# Resolution



Full-width half-max (FWHM) =  $1.22 \frac{\lambda}{D}$  radians  
 $\lambda$  = wavelength,  $D$  = aperture diameter

# Camera/Telescope Details

Aperture:  $D = 10 \text{ inches} = 254 \text{ mm}$

Wavelength:  $\lambda \approx 5080 \text{ Angstroms} = 508 \text{ nm} = 0.508 \text{ microns} = 0.000508 \text{ mm}$

FWHM =  $1.22 \frac{\lambda}{D} = 1.22 \frac{0.000508}{254} = 0.00000244 \text{ radians} = 0.503 \text{ arcseconds}$

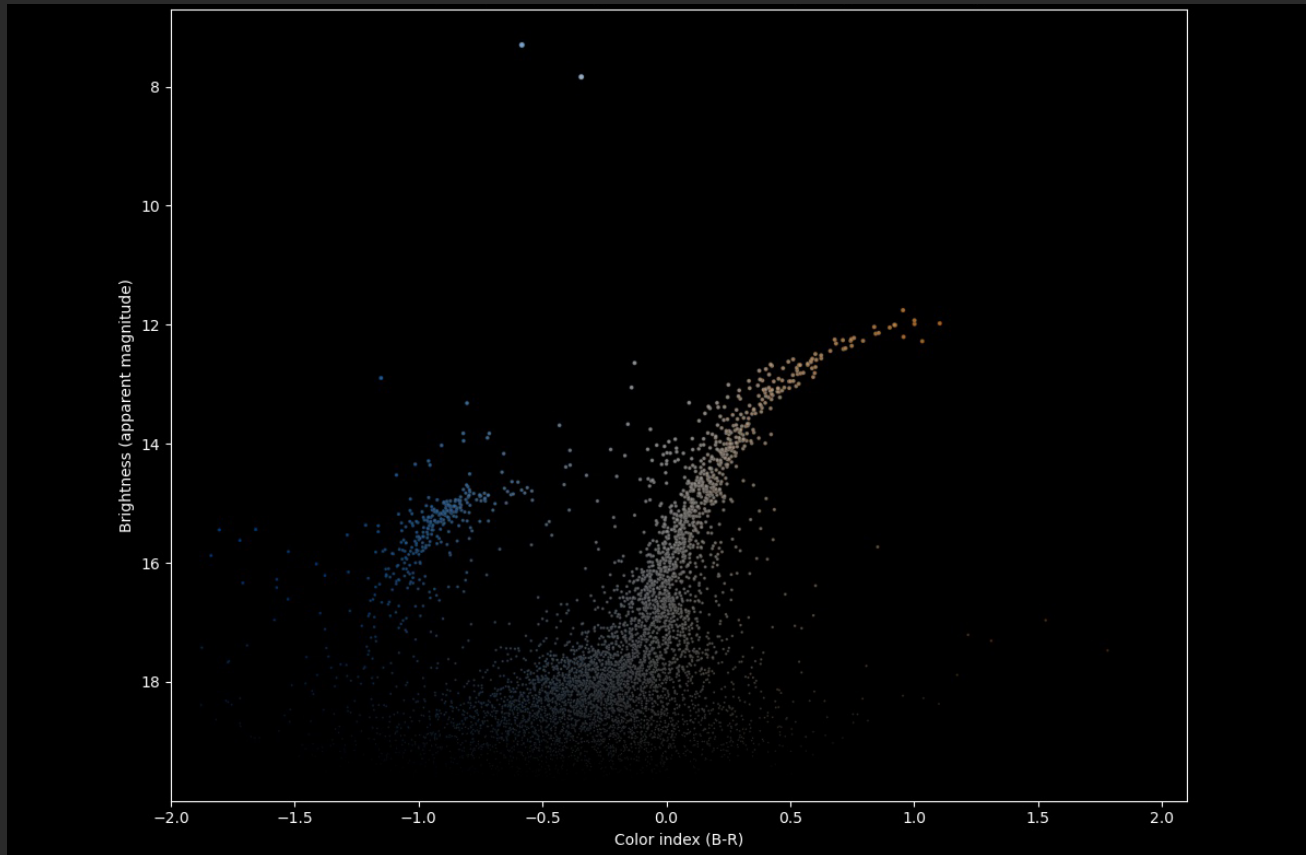
Focal Length:  $f = 90 \text{ inches} = 2286 \text{ mm}$

FWHM in microns =  $0.00000244 \text{ radians} \times 2286 \text{ mm} \times 1000 \text{ microns/mm} = 5.58 \text{ microns}$

Pixel Size:  $6.4 \text{ microns/pixel}$

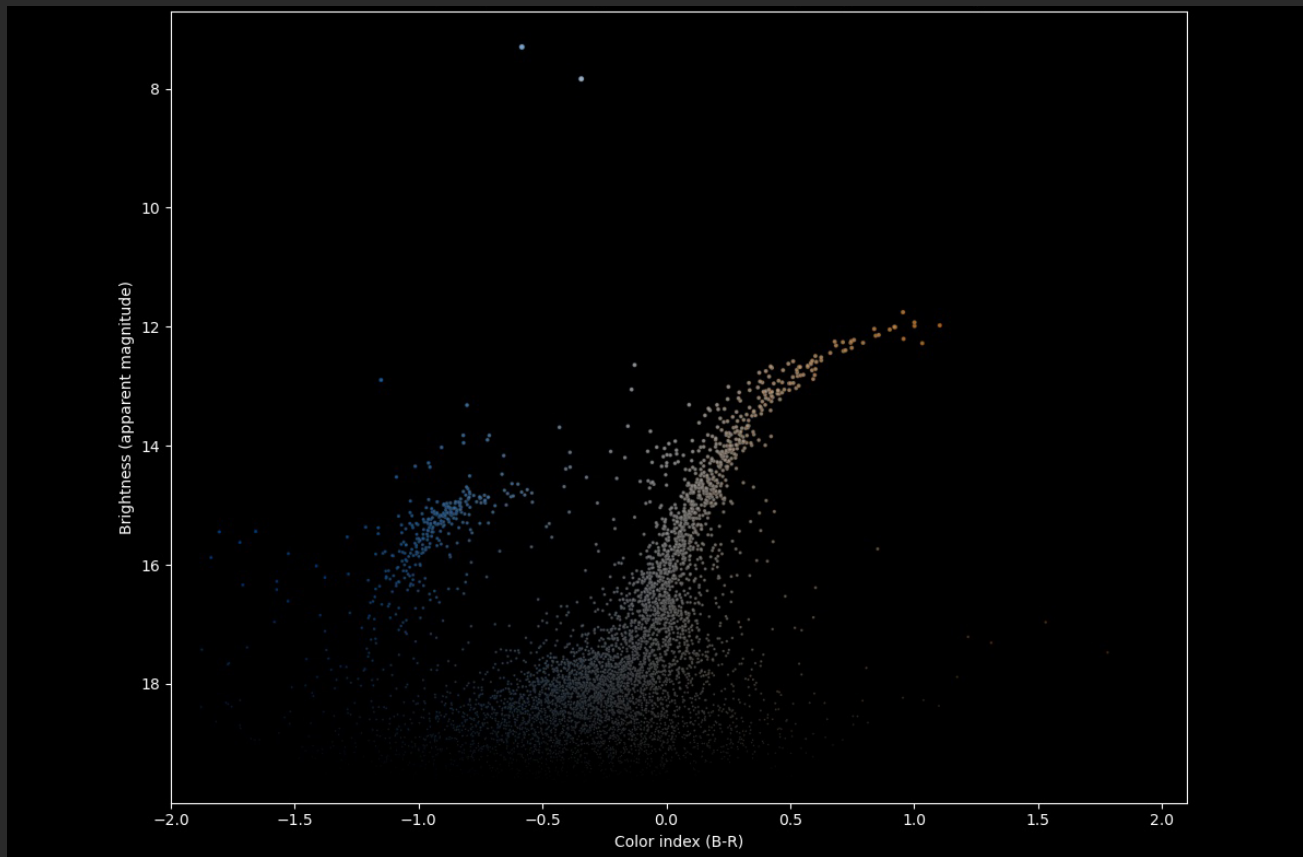
FWHM in pixels:  $5.58/6.4 = 0.87 \text{ pixels}$

# Measuring Distances: RR-Lyrae and M13



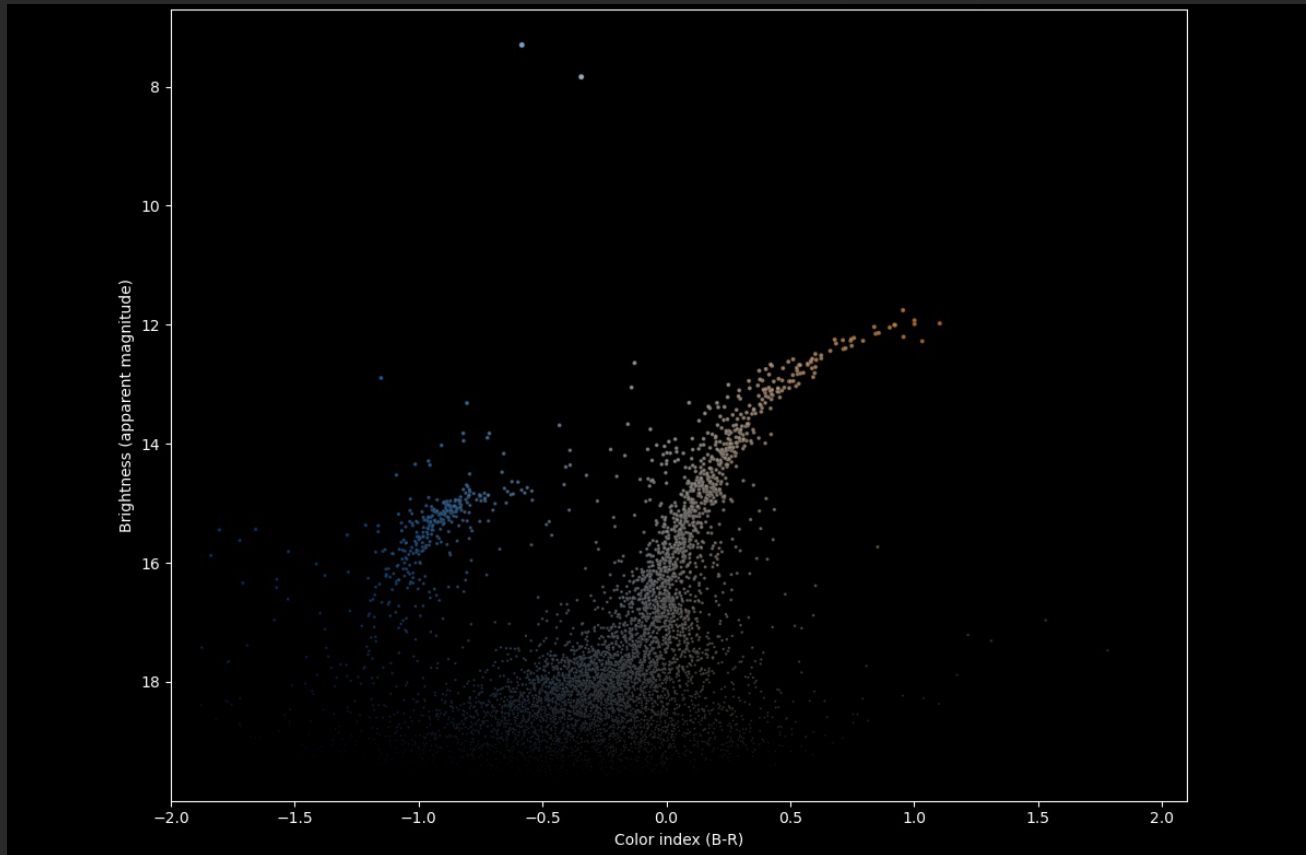
Assuming that the distance to RR-Lyrae is known, we can overlay an image of RR-Lyrae on an image of M13 to estimate the distance to the cluster.

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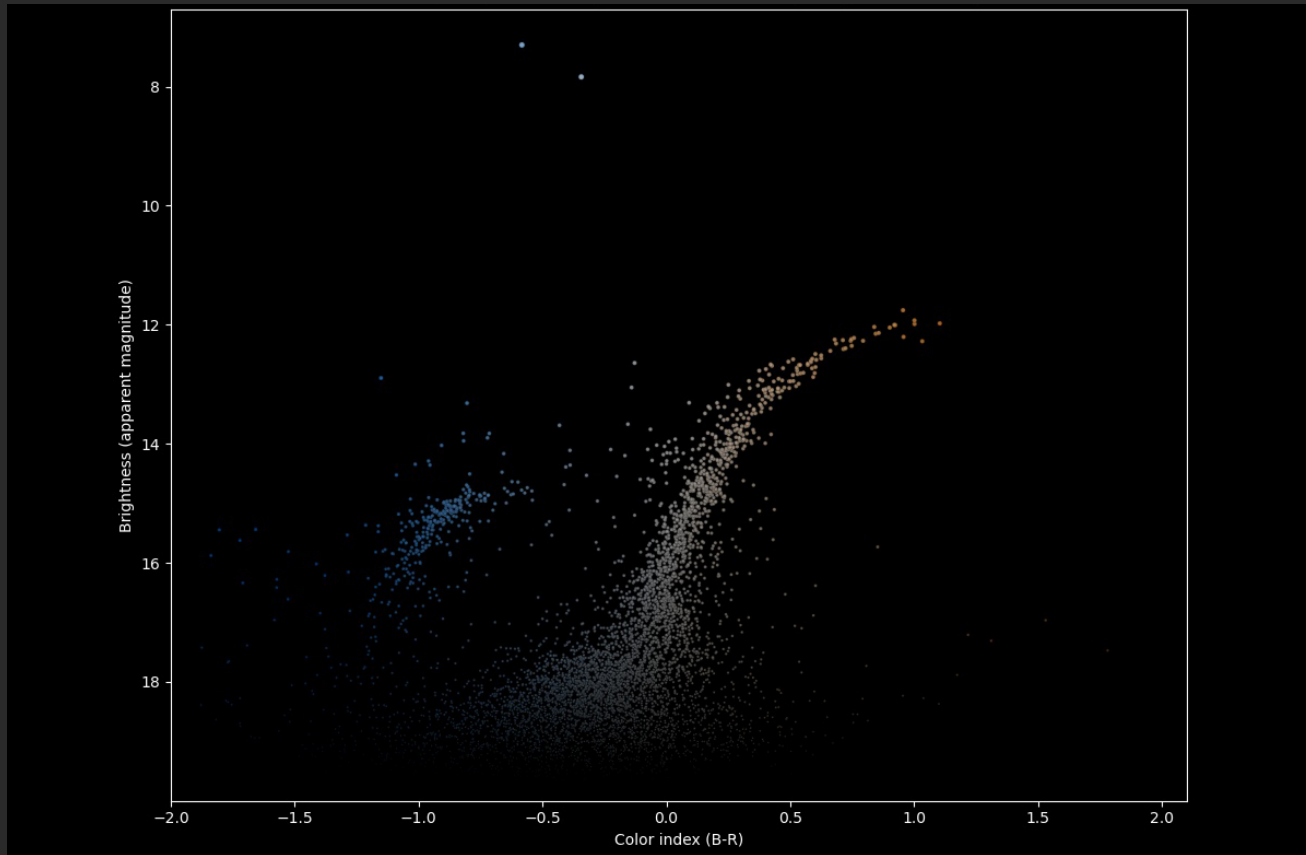
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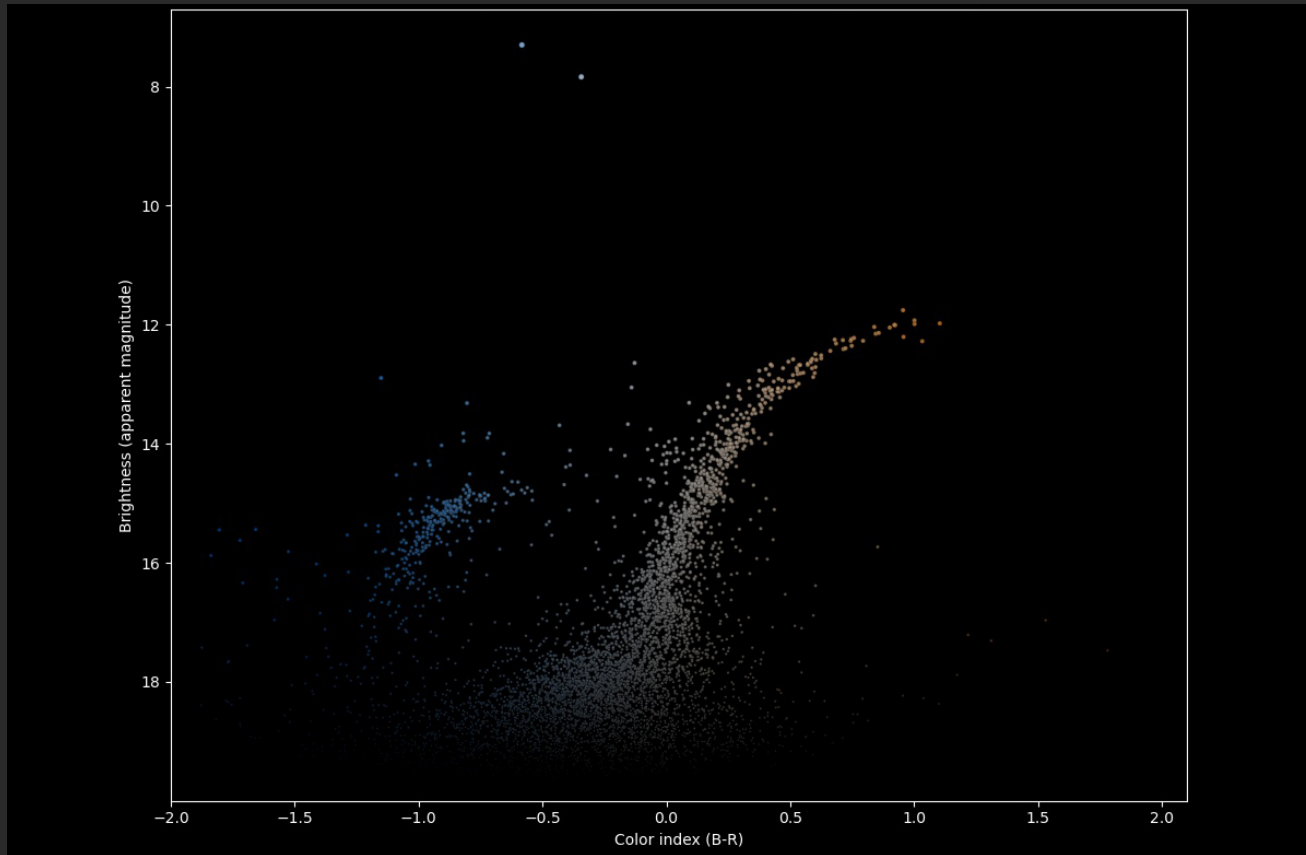
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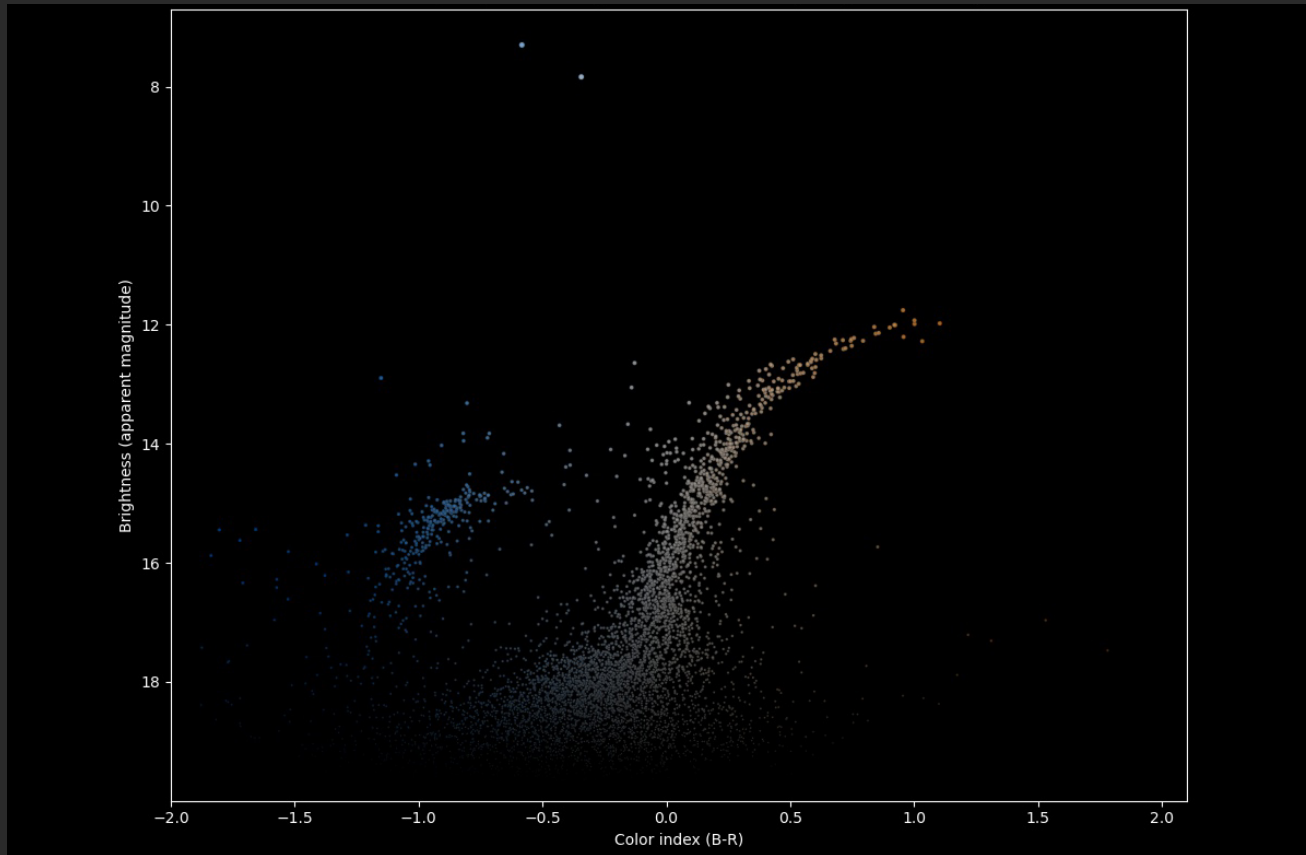
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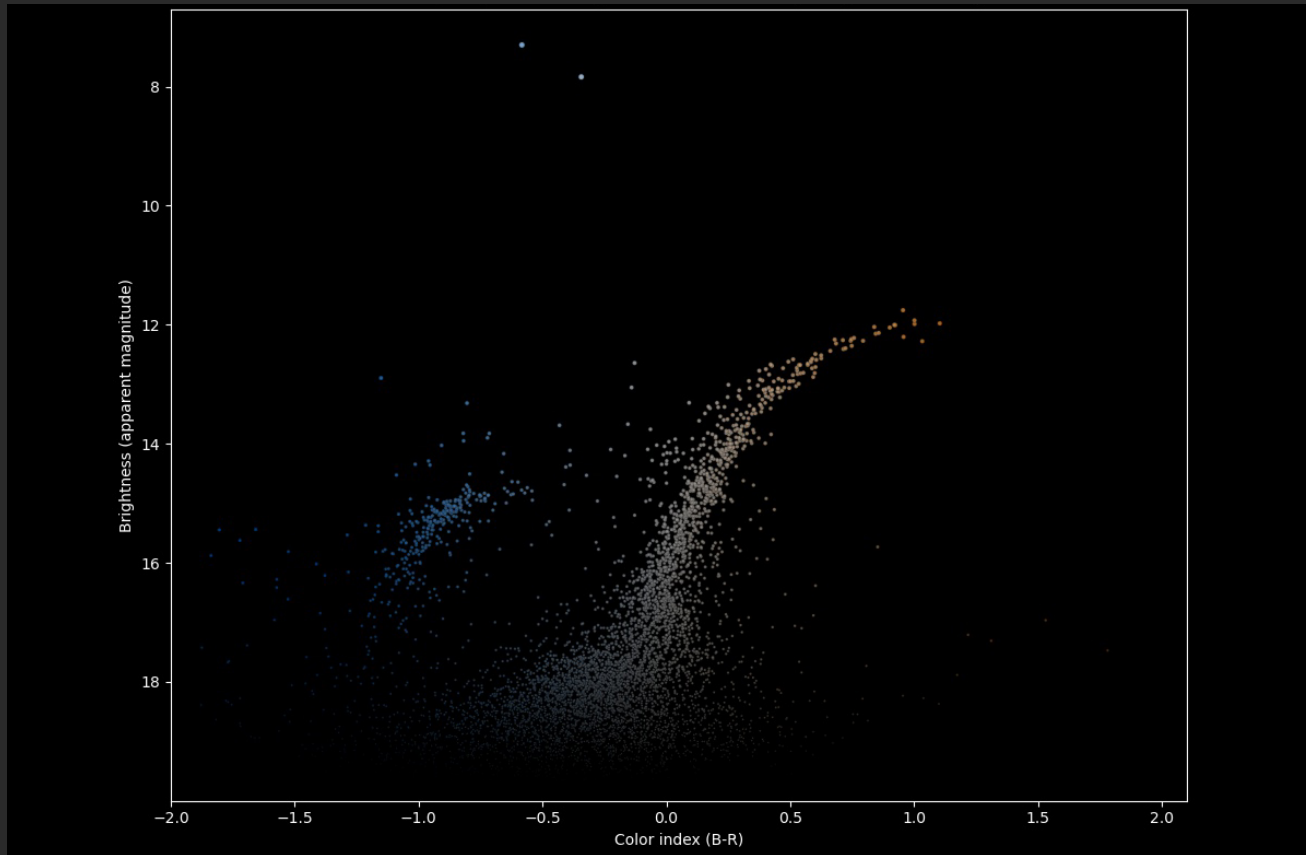
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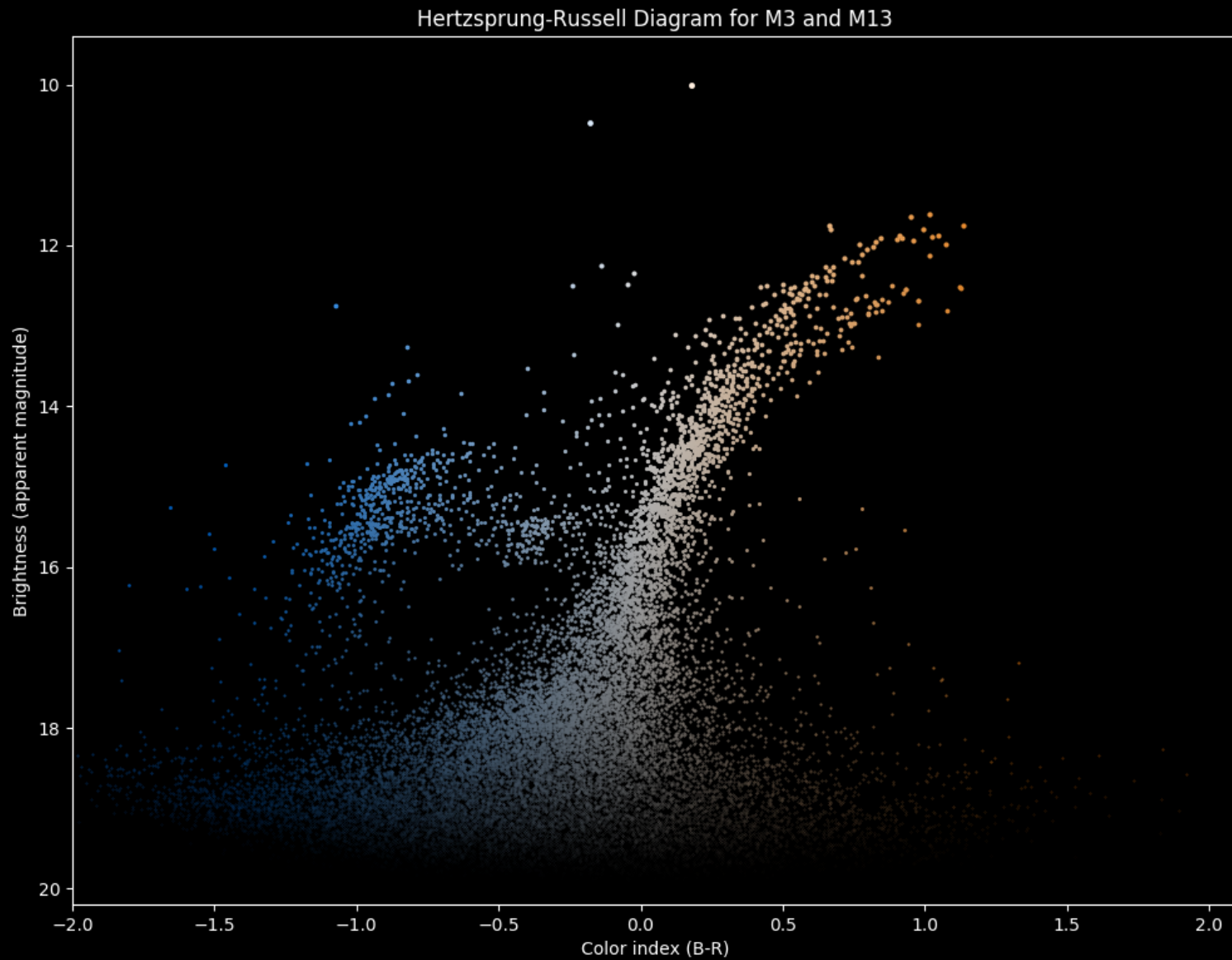
# M3/M13 Comparison: Here's M3



# M3/M13 Comparison: Here's M13



# M3/M13 Comparison



# M3/M13 Comparison

Click [here](#) to download the Python code.

The fits files can be accessed here:

[https://vanderbei.princeton.edu/FRS\\_131/python/fits\\_files/m3-RGB.fit](https://vanderbei.princeton.edu/FRS_131/python/fits_files/m3-RGB.fit)

[https://vanderbei.princeton.edu/FRS\\_131/python/fits\\_files/m13-RGB.fit](https://vanderbei.princeton.edu/FRS_131/python/fits_files/m13-RGB.fit)

Here's the output from Python:

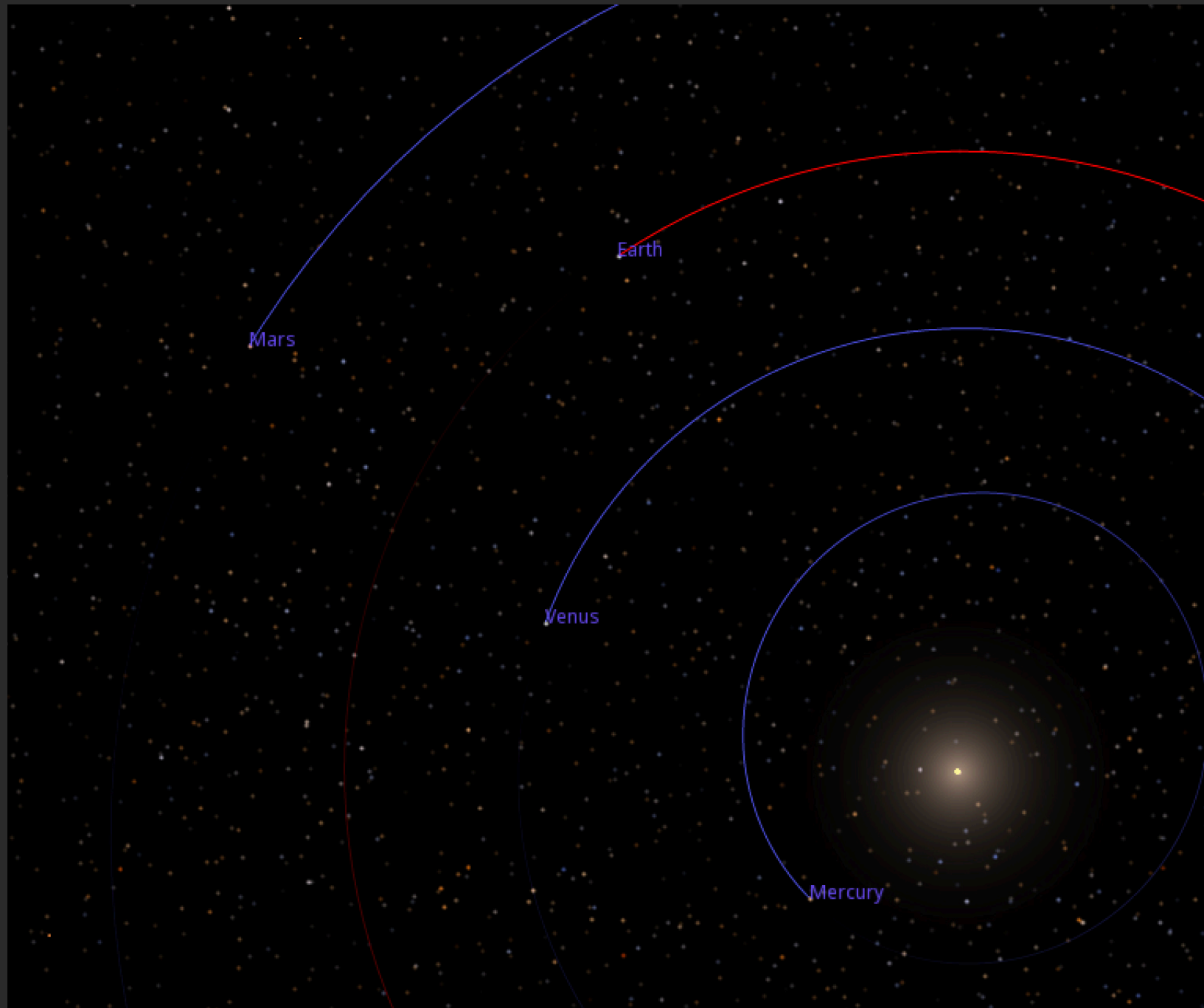
```
difference in brightness is about 0.8 magnitude  
difference in flux = 10^(0.8/2.5) = 2.2909  
relative distance factor = sqrt(flux) = 1.5136
```

From Wikipedia, we see that the true distances are:

```
M3 = 10.4 kpc and M13 = 6.8 kpc  
true distance ratio = 10.4/6.8 = 1.5294
```

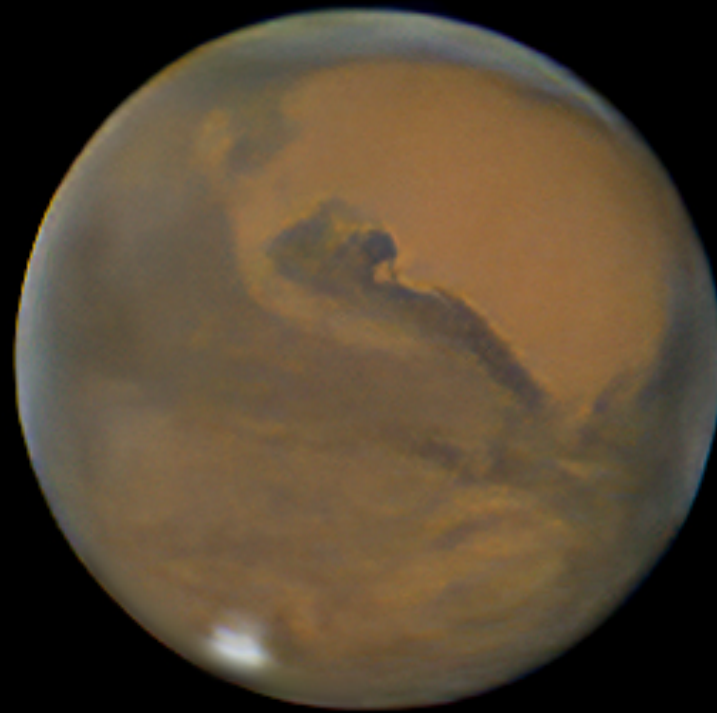
# A Brief Step Back Toward Home

# Earth "Passed" Mars in Oct. 2020



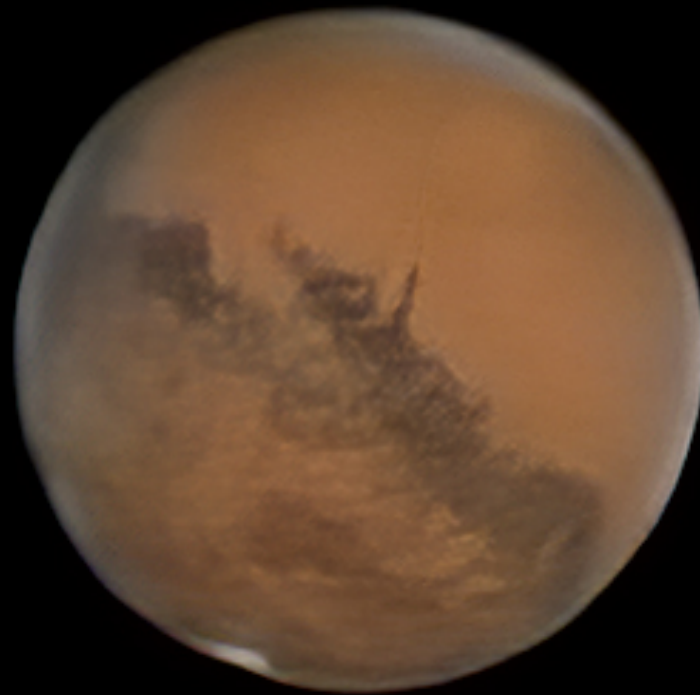
Mars

Oct. 6, 2020



Mars

Oct. 18, 2020



Questions?